

# MULTISPECTRAL COMPOSITES AND NEXT-GENERATION ADVANCED SATELLITE IMAGERS

EMILY BERNDT<sup>12</sup>, NICHOLAS ELMER<sup>13</sup>,  
LORI SCHULTZ<sup>13</sup>, AND ANDREW MOLTHAN<sup>12</sup>

<sup>1</sup>SHORT-TERM PREDICTION RESEARCH AND TRANSITION CENTER

<sup>2</sup>NASA MARSHALL SPACE FLIGHT CENTER

<sup>3</sup>UNIVERSITY OF ALABAMA IN HUNTSVILLE

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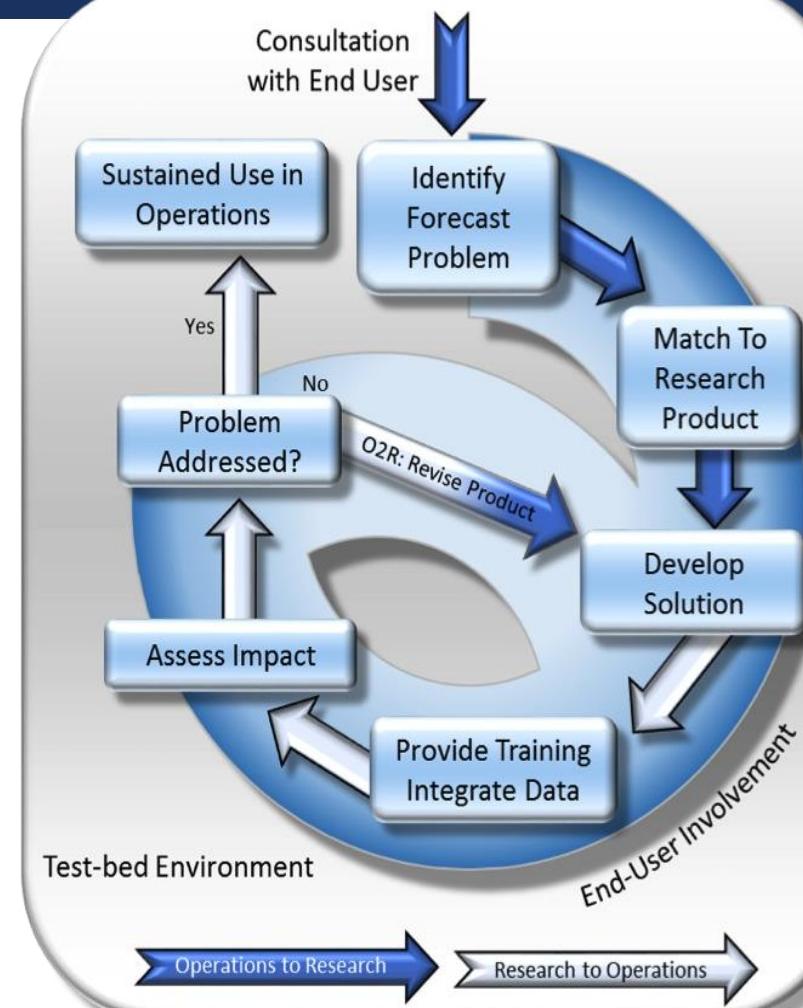


**SPORT**

Short-term Prediction Research and Transition Center

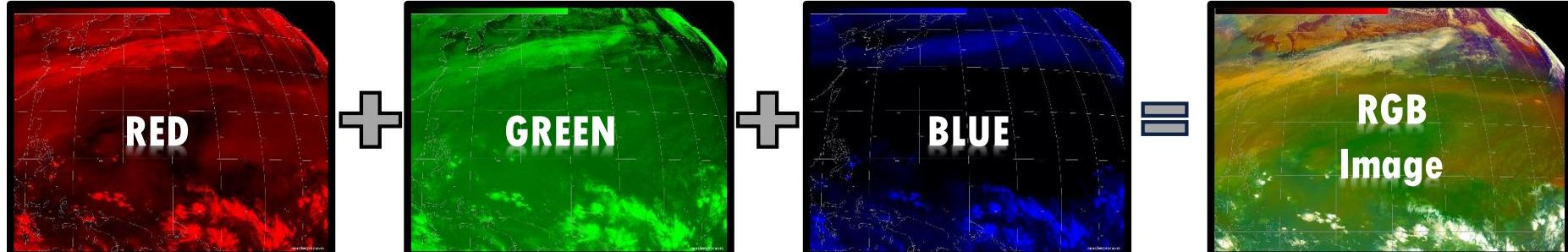
# NASA/SPORT MISSION AND PARADIGM

- **Apply satellite measurement systems and unique Earth science research to improve the accuracy of short-term weather prediction at the regional and local scale**
- Bridge the “Valley of Death” between research and operations
- Can’t just “throw data over the fence”
  - Maintain collaborative partnerships with end users via help of local “SPoRT” advocates
  - Integrate product into user decision support tools for use with existing data
  - Create forecaster training on product utility
  - Perform targeted product assessments to determine operational value
- Concept has been used to successfully transition a variety of satellite datasets to operational users for nearly 10 years



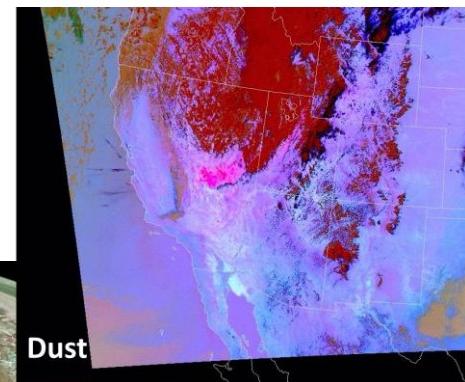
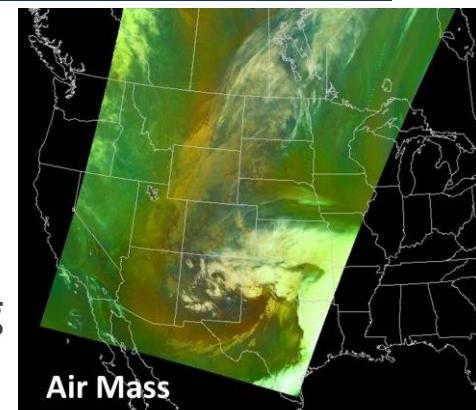
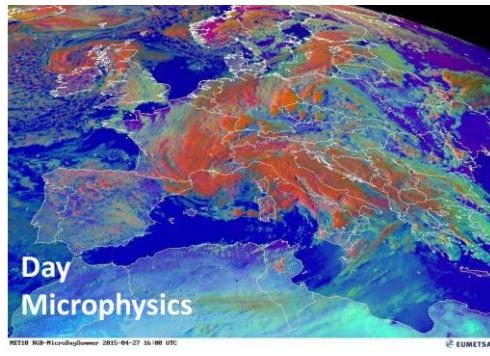
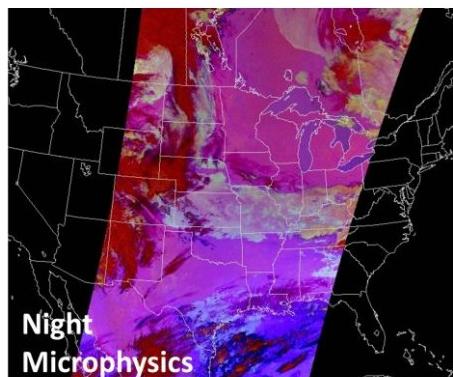
# INTRODUCTION

- EUMETSAT began creating Multispectral Composite (i.e. RGB) imagery in the early 2000s with the launch of Meteosat Second Generation Spinning EnhancedVisible and Infrared Imager (SEVIRI)
- RGB imagery is the use of single channels or channel differences combined into each of the red, green, and blue color components, resulting in a false-color composite related to multiple atmospheric and land-surface features
- RGB products are qualitative in nature are designed to enhance a specific phenomena such as low clouds and fog, dust, convection, air mass characteristics, or volcanic ash
- SPoRT has invested research in creating consistent Multispectral Composite (i.e. RGB) imagery across different sensors onboard polar-orbiting and geostationary satellites (Elmer et al. 2016)



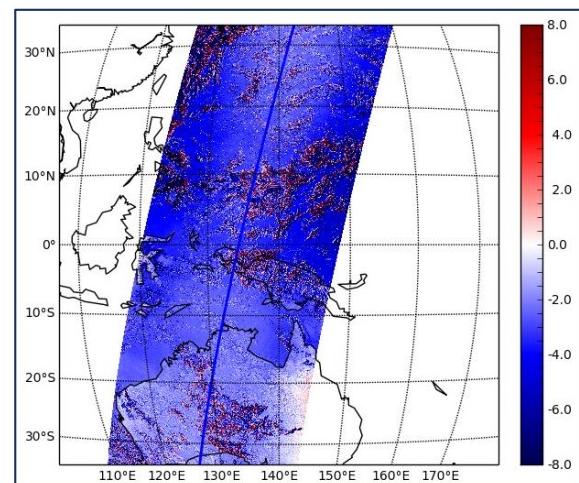
# INTRODUCTION

- EUMETSAT developed a set of best practices to identify a minimum set of multispectral composites based on MSG SEVIRI
- The best practices recommend adjusting the RGB recipe when creating RGBs with instruments such as the Moderate Resolution Imaging Spectroradiometer (MODIS) due to differing spectral and absorption characteristics across sensors
- Differences in band central wavelength, bandwidth, response functions and atmospheric absorption between sensors can result in inconsistencies in an RGB composites from sensor to sensor



# METHODOLOGY

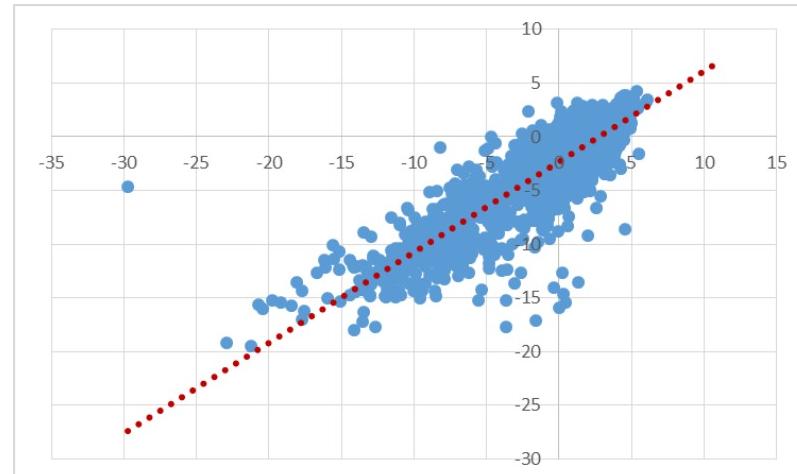
- For the purpose of comparing RGB composites derived from different sensors, Elmer et al. (2016) developed an offset correction to account for differences in band spectral response
- The offset correction and limb correction was applied to VIIRS to create a reference sensor or SEVIRI proxy
- Brightness temperatures for the SEVIRI proxy and AHI were compared over cloud-free ocean scenes at shared nadir points and less than 10 minutes apart similar to Cao et al. (2014) and Wu et al. (2012)
- Since atmospheric and surface changes are negligible at shared nadir points (cloud-free ocean scenes), any differences in brightness temperature between the two sensors can be attributed solely to differences in band spectral response



*Difference between AHI and SEVIRI proxy for comparable 3.9  $\mu\text{m}$  bands*

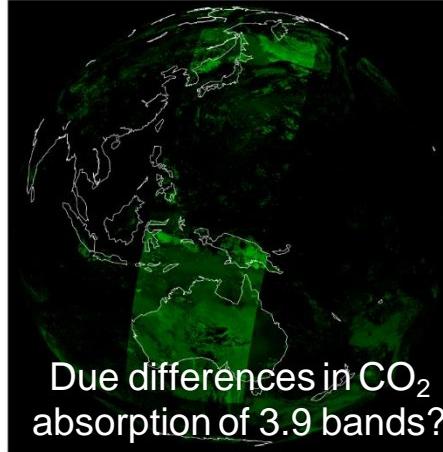
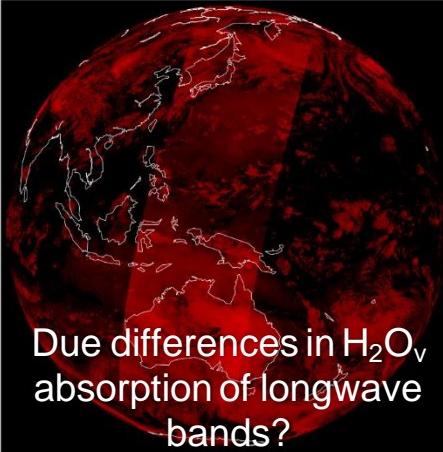
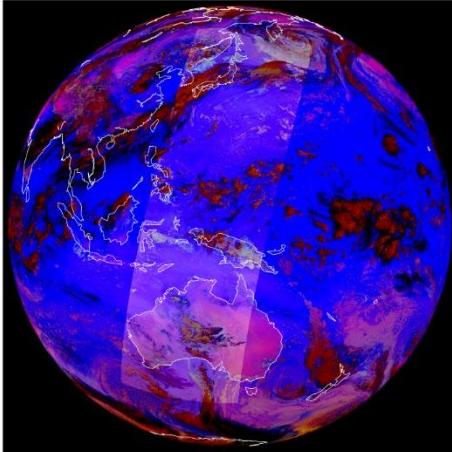
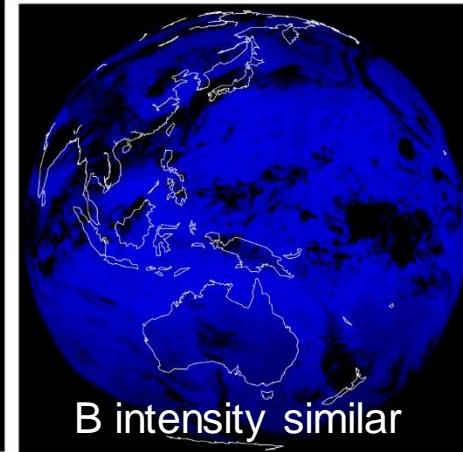
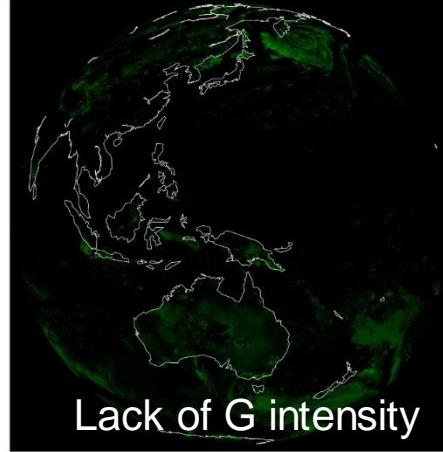
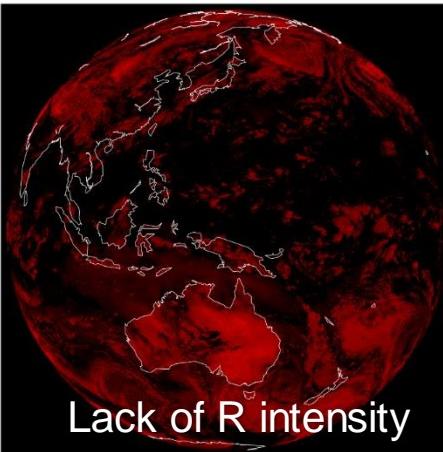
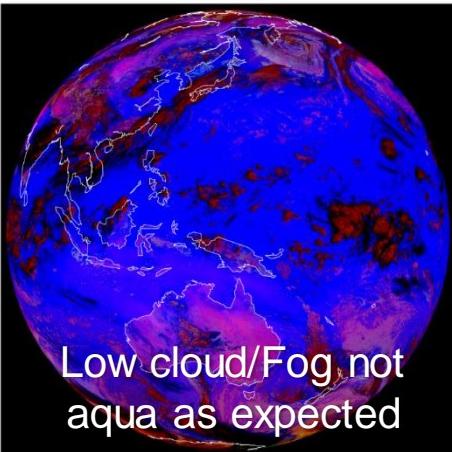
# METHODOLOGY

- The relationship between measured brightness temperature of the SEVIRI proxy and AHI was compared through linear regression
- The linear regression coefficients were then applied to the EUMETSAT RGB component minimum and maximum values similar to Shimzu (2015) to determine RGB recipe adjustments
- By adjusting the recipe to account for differences in spectral characteristics and response the result is RGB imagery consistent with legacy EUMETSAT images and interpretation



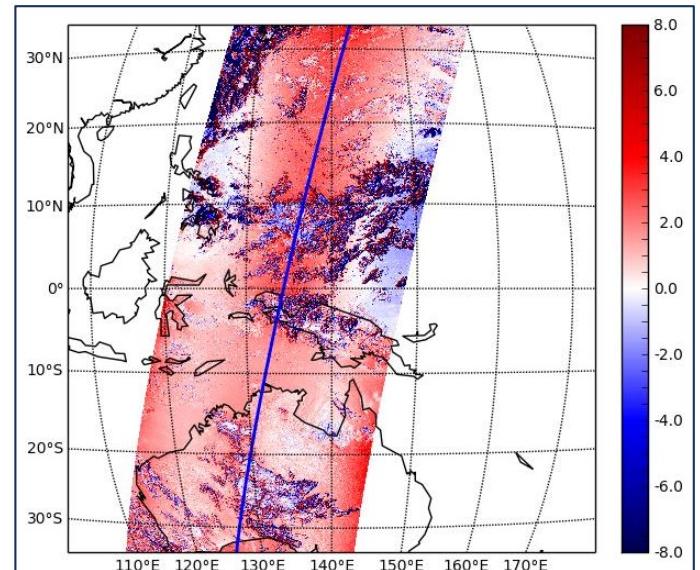
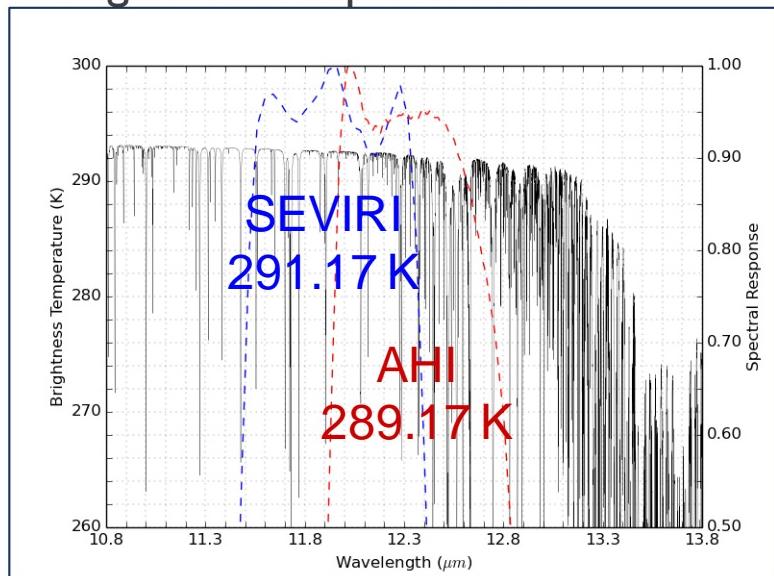
# NIGHT-TIME MICROPHYSICS RGB

- AHI RGB coloring and component intensities differ from the SEVIRI proxy overlay



# COMPARABLE 12.0 MICRON BANDS

- Case study data show SEVIRI proxy 12.0  $\mu\text{m}$  is 2 to 4 K warmer than AHI 12.4  $\mu\text{m}$
- LBLRTM simulations confirmed the 2 K offset
- Despite a narrower range than SEVIRI, AHI 12.4  $\mu\text{m}$  shifted to a region with more water vapor absorption and results in cooler measured brightness temperatures

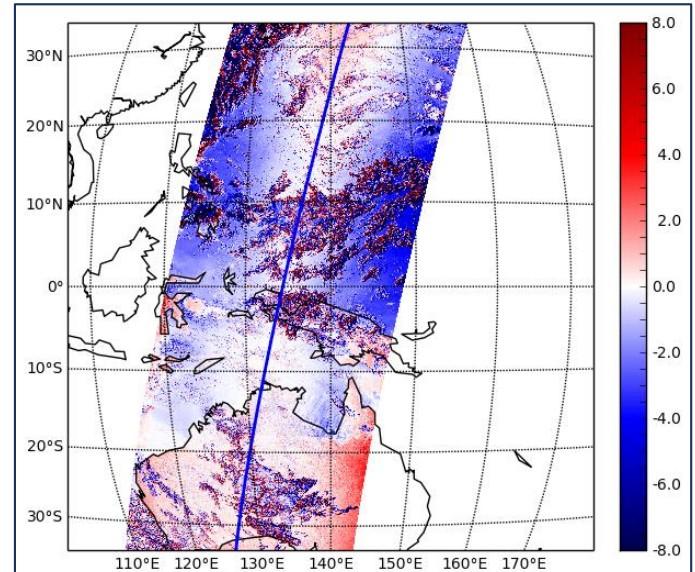
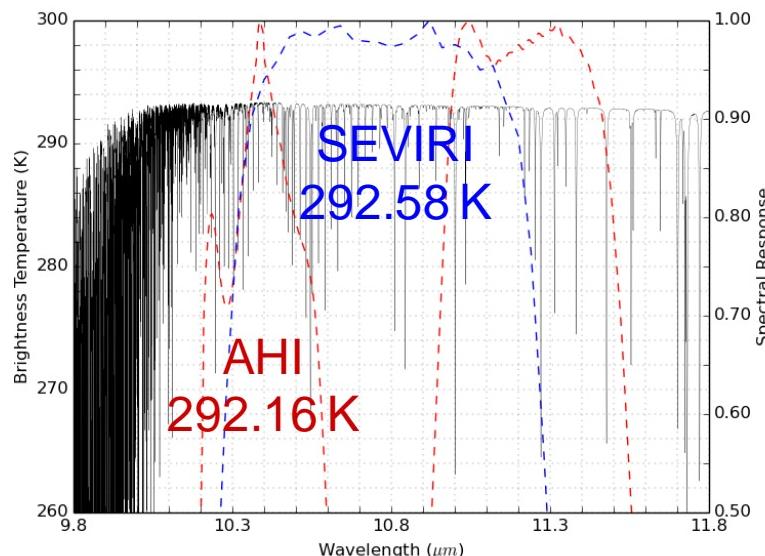


*Difference between AHI and SEVIRI proxy for comparable 12.0  $\mu\text{m}$  bands*

*LBLRTM simulation using the U.S. Standard atmospheric profile, nadir viewing (100 km to surface) brightness temperature (Black), SEVIRI (blue) and AHI (red) spectral responses*

# COMPARABLE 10.8 MICRON BANDS

- Case study data show SEVIRI proxy 10.8  $\mu\text{m}$  is nearly equal, slightly warmer than AHI 10.4  $\mu\text{m}$
- LBLRTM simulations confirmed the slight offset
- Brightness temperatures are similar in the atmospheric window despite shifted AHI central wavelength and narrow spectral range

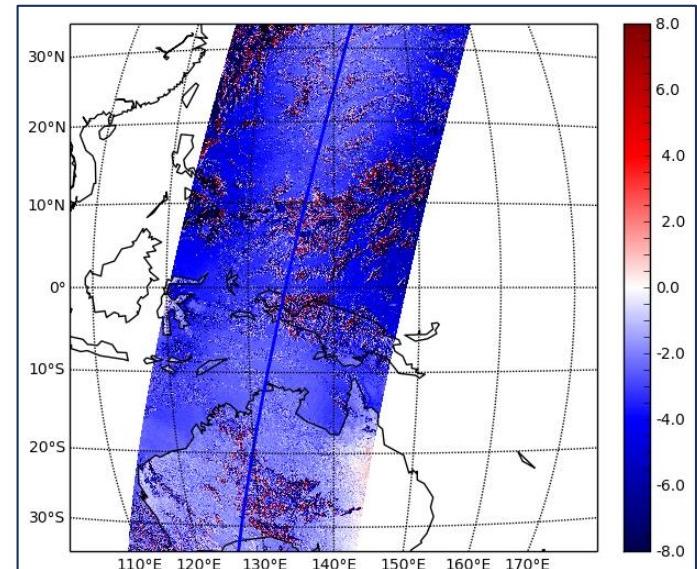
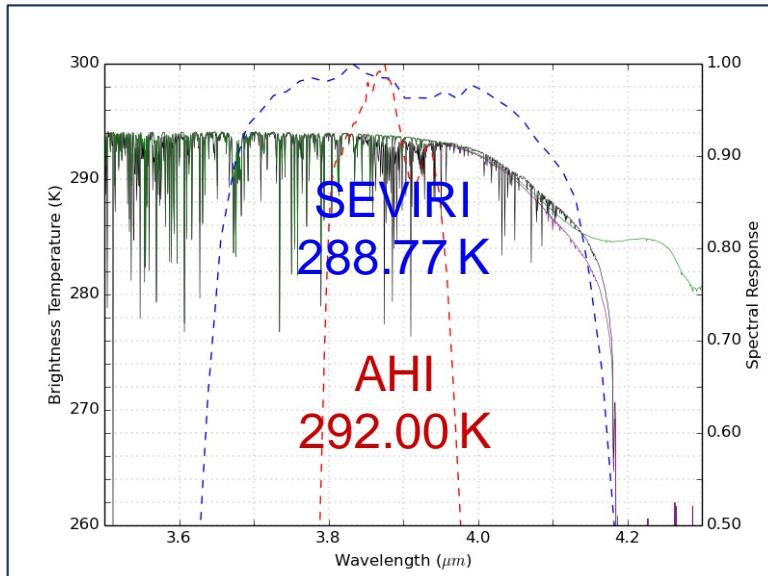


*Difference between AHI and SEVIRI proxy for comparable 10.8  $\mu\text{m}$  bands*

*LBLRTM simulation using the U.S. Standard atmospheric profile, nadir viewing (100 km to surface) brightness temperature (Black), SEVIRI (blue) and AHI (red) spectral responses*

# COMPARABLE 3.9 MICRON BANDS

- Case study data show SEVIRI proxy 3.9  $\mu\text{m}$  is 2 to 4 K cooler than AHI 3.9  $\mu\text{m}$
- LBLRTM simulations confirmed the offset
- AHI 3.9  $\mu\text{m}$  brightness temperature are warmer since the band is not influenced by carbon dioxide absorption due to narrow spectral range



*Difference between AHI and SEVIRI proxy for comparable 3.9  $\mu\text{m}$  bands*

*LBLRTM simulation using the U.S. Standard atmospheric profile, nadir viewing (100 km to surface) brightness temperature for 7 gases (Black),  $\text{H}_2\text{O}_v$  and  $\text{CO}_2$  (Purple),  $\text{H}_2\text{O}_v$  only (Green), SEVIRI (blue) and AHI (red) spectral responses*

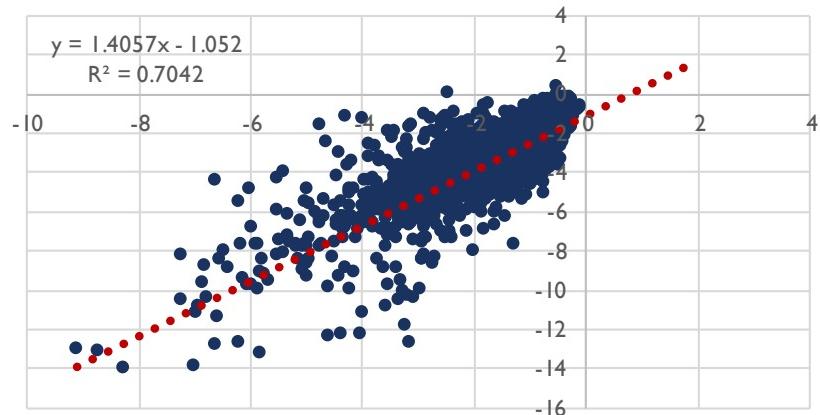
# LINEAR REGRESSION: RED COMPONENT ADJUSTMENT

- Scatter plots and linear regression for two case studies used to determine recipe adjustments

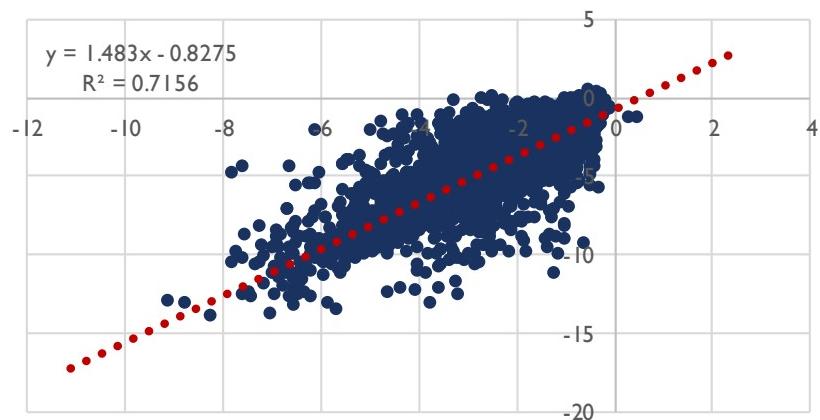
Night-time Microphysics RGB				
Component	Band/Band Difference	Min	Max	Gamma
R	12.0 – 10.8	-4 K	2 K	1.0
G	10.8 – 3.9	0 K	10 K	1.0
B	10.8	243 K	293 K	1.0

- For Case 1, new range = -6.7 to 1.8
- For Case 2, new range = -6.7 to 2.1
- JMA adjustment = -6.7 to 2.6
- Results suggest adjustment:
  - Is not seasonally dependent
  - Is on par with JMA adjustments which were based on simulated data

Case 1: 30 Aug 2015 SEVIRI vs AHI



Case 2: 14 Dec 2016 SEVIRI vs AHI



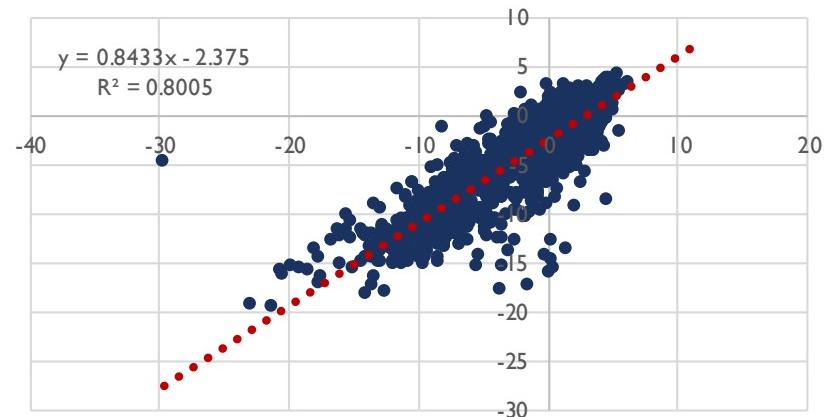
# LINEAR REGRESSION: GREEN COMPONENT ADJUSTMENT

- Scatter plots and linear regression for two case studies used to determine recipe adjustments

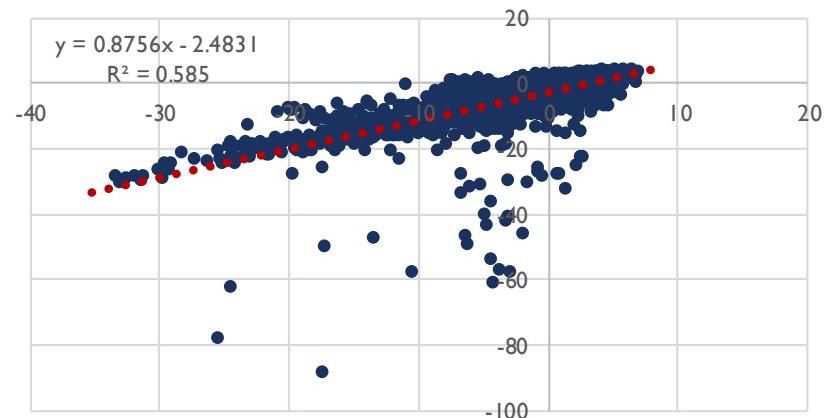
Night-time Microphysics RGB				
Component	Band/Band Difference	Min	Max	Gamma
R	12.0 – 10.8	-4 K	2 K	1.0
G	10.8 – 3.9	0 K	10 K	1.0
B	10.8	243 K	293 K	1.0

- For Case 1, new range = -2.4 to 6.1
- For Case 2, new range = -2.4 to 6.4
- JMA adjustment = -3.1 to 5.2
- Results suggest adjustment:
  - Is not seasonally dependent
  - Is on par with JMA adjustments which were based on simulated data

Case 1: 30 Aug 2015 SEVIRI vs AHI



Case 2: 14 Dec 2016 SEVIRI vs AHI



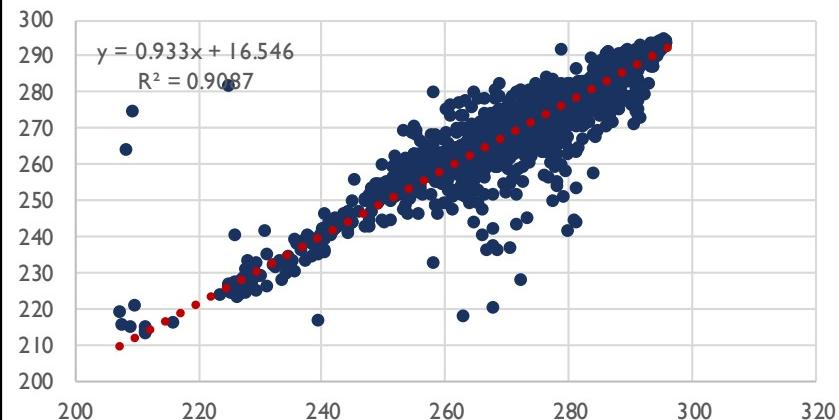
# LINEAR REGRESSION: BLUE COMPONENT ADJUSTMENT

- Scatter plots and linear regression for two case studies used to determine recipe adjustments

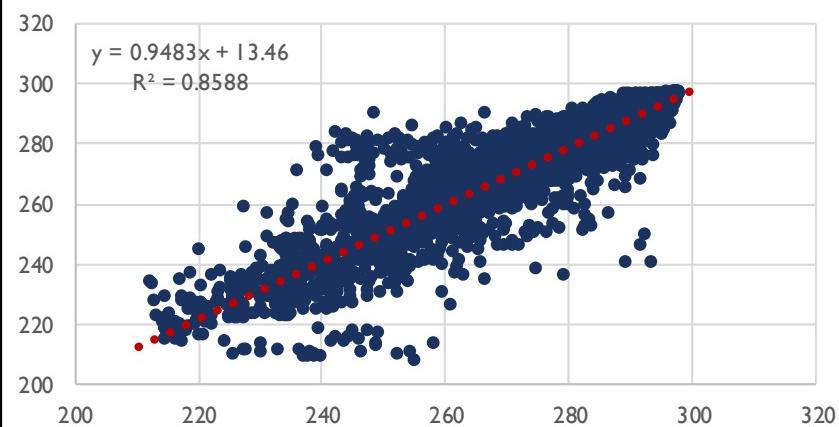
Night-time Microphysics RGB				
Component	Band/Band Difference	Min	Max	Gamma
R	12.0 – 10.8	-4 K	2 K	1.0
G	10.8 – 3.9	0 K	10 K	1.0
B	10.8	243 K	293 K	1.0

- For Case 1, new range = -244.2 to 292.1
- For Case 2, new range = 243.9 to 291.3
- JMA adjustment = 243.6 to 292.6
- Results suggest adjustment:
  - Is not seasonally dependent
  - Is on par with JMA adjustments which were based on simulated data

Case 1: 30 Aug 2015 SEVIRI vs AHI



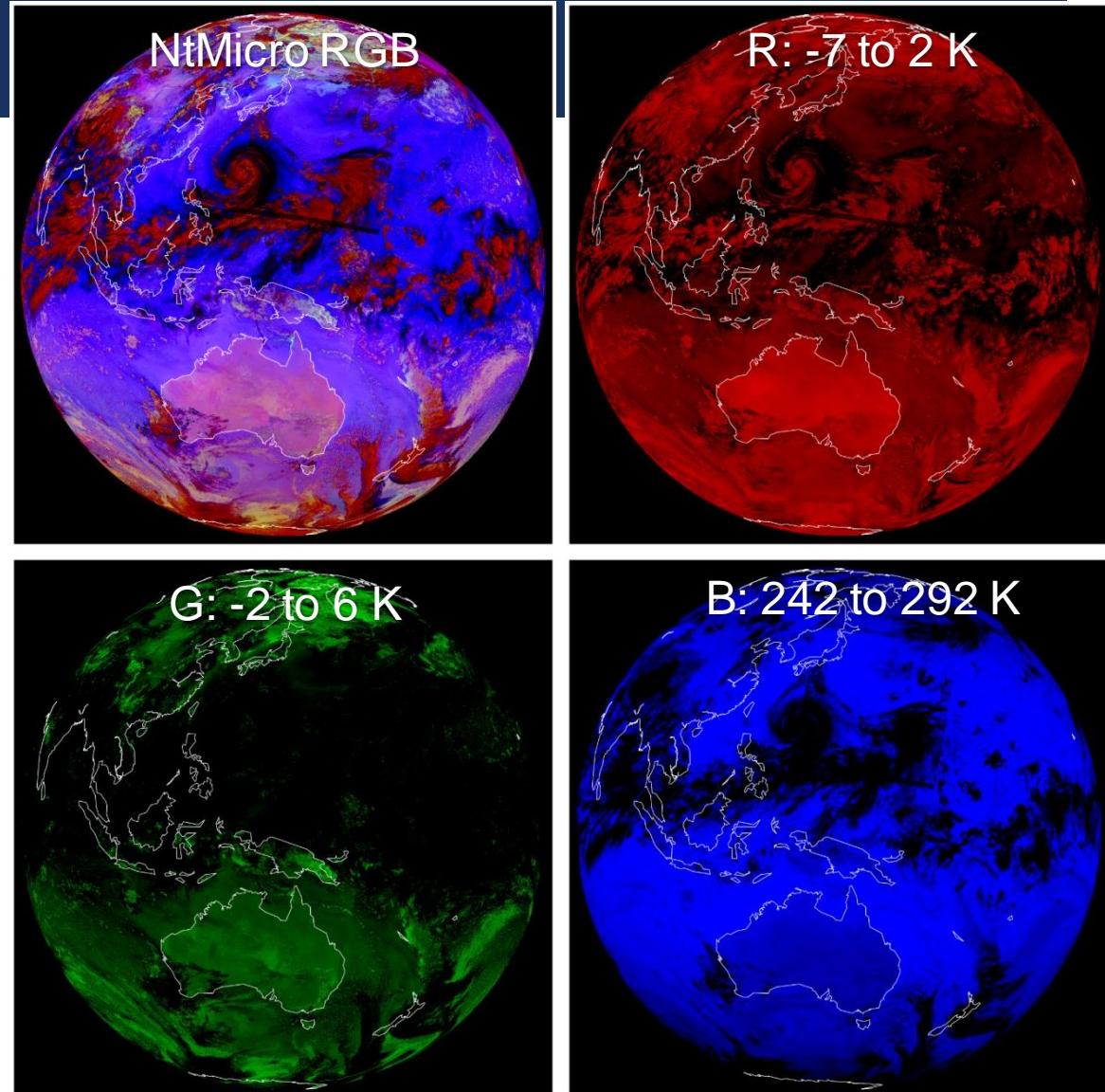
Case 2: 14 Dec 2016 SEVIRI vs AHI



# RECIPE ADJUSTMENT

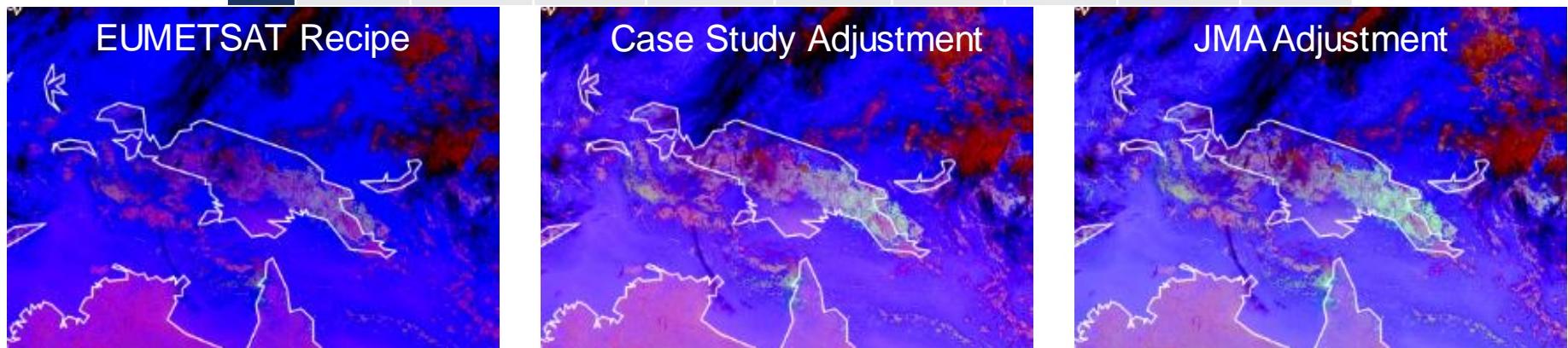
- With the new adjustment RGB colors and component intensity are similar to the SEVIRI proxy overlay (SEVIRI overlay barely visible!)
- Adjustment improves the lack of red and green intensities noted earlier
- RGB colors are consistent with legacy interpretation and training

\*note: applied adjustment to case other than which it was derived



# RECIPE ADJUSTMENT

	EUMETSAT			Adjustment by Case Studies			Adjustment by JMA		
	Min (K)	Max (K)	Gamma	Min (K)	Max (K)	Gamma	Min (K)	Max (K)	Gamma
R	-4	2	1	-7	2	1	-6.7	2.6	1
G	0	10	1	-2	6	1	-3.1	5.2	1
B	243	293	1	244	292	1	243.6	292.6	1



- Both adjustments improve the aqua coloring of the low cloud/fog features
- Slight color differences between case study and JMA adjustments attributed to the ability to account for instrument bias, noise, and full atmospheric absorption when using real data
- Since the adjustment is similar to what JMA derived for AHI, this research verifies work done by JMA and demonstrates a methodology to determine recipe adjustments for RGB imagery derived with GOES-R and GOES-S in the future

# SUMMARY

- SPoRT has invested research in creating consistent Multispectral Composite (i.e. RGB) imagery across different sensors onboard polar-orbiting and geostationary satellites
- Differences in band central wavelength, bandwidth, response functions and atmospheric absorption between sensors can result in inconsistencies in an RGB composites from sensor to sensor
- This research presented a methodology to adjust RGB recipes to account for differing spectral characteristics between sensors through case study analysis including:
  - Comparison with a reference radiometer
  - Linear regression
- This methodology was applied to the Night-time Microphysics RGB derived from AHI and
  - Resulted in adjustments similar to JMA
  - Demonstrates a methodology to adjust RGB recipes for GOES-R and GOES-S
- Although not shown the methodology was applied to adjust recipes for the Dust, Ash, and 24-hour Microphysics RGB composites
- The methodology can be applied to visible and water vapor bands to derive recipe adjustments for additional EUMETSAT RGB composites

# REFERENCES

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QUESTIONS?



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**Short-term Prediction Research and Transition Center**

<http://weather.msfc.nasa.gov/sport/>

<https://nasasport.wordpress.com/>

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